

Partie VIII

Bibliographie

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Instabilités et sources locales de turbulence dans les disques d'accrétion

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Résumé

Le problème du transport de moment cinétique dans les disques d'accrétion astrophysiques fait l'objet d'un vif débat depuis maintenant une trentaine d'années. On propose ici une étude de ce transport en considérant plusieurs instabilités pouvant conduire à de la turbulence développée dans les disques. La première instabilité proposée est l'instabilité hydrodynamique sous critique de cisaillement. Une étude numérique de cette instabilité nous permet de montrer que, bien que non détectable dans les simulations, le transport obtenu sera très probablement bien trop faible pour expliquer les observations actuelles. Nous étudions ensuite l'instabilité stratorotationnelle et nous montrons par un développement analytique que les contraintes sur les conditions aux limites empêchent la formation de cette instabilité dans un vrai disque. Enfin, nous nous intéressons à l'instabilité magnéto-rotationnelle en présence d'effets non idéaux (viscosité et résistivité magnétique). Nous montrons alors que ces effets non idéaux, bien que faibles a priori, ont un fort impact sur l'efficacité de la turbulence finalement engendrée. Ce dernier point montre le rôle que peut avoir la microphysique dans le processus de saturation, et la nécessité d'avoir une modélisation de ces effets afin d'obtenir un modèle de transport cohérent dans les disques d'accrétion.

Mots clé : Disque d'accrétion, méthodes numériques, théorie de la turbulence, MHD

Abstract

Angular momentum transport in accretion disks has been a highly controversial debate for 30 years now. We present here a study of this transport, considering some instabilities that can lead to developed turbulence in disks. The first instability considered is the subcritical hydrodynamic instability. Our numerical study shows that, although not directly observed in simulations, this instability should lead to a very weak turbulent transport, and is probably not relevant to explain disks observations. We then consider the stratorotational instability, and we show using an analytical approach that it requires very specific boundary conditions, which prevent this instability from appearing in disks. Finally, we study the magnetorotational instability, including non ideal MHD effect (resistivity and viscosity). We show numerically that these effects, although small compared to ideal MHD terms, can have a strong impact on the turbulence efficiency. This point emphasizes the role of microphysic effects in the saturation process, and the necessity to modelize these phenomena to get a self-consistent accretion disk transport model.

Keywords : Accretion disc, numerical methods, turbulence theory, MHD